

# Mag - Tutorial 2 - Reversal Dynamo Case

## Tutorial 2: Reversal Dynamo Case

In this example, we produce a magnetic field reversal using MAG. The input parameter in the source directory for this case is `~/src/par.Rev`. There is no longitudinal symmetry in this case, so when you compile MAG, use `param32s1.f` linking to `param.f`. The Ekman number is  $E=0.02$ , the Prandtl number is  $Pr=1$  and the magnetic Prandtl number is  $Pm=10$ . The Rayleigh number is  $Ra=12000$ .

Link `param32s1.f` to `param.f`:

```
$ ln -sf param32s1.f param.f
```

Compile the program with:

```
$ make
$ mv magx magx32s4
```

Background execute using `par.Rev` as the input file and `p.rev` as the output file (Note: The output file's extension is specified in the par-file, e.g., here `par.Rev` specifies that the outfile should be named `*filename*.rev`):

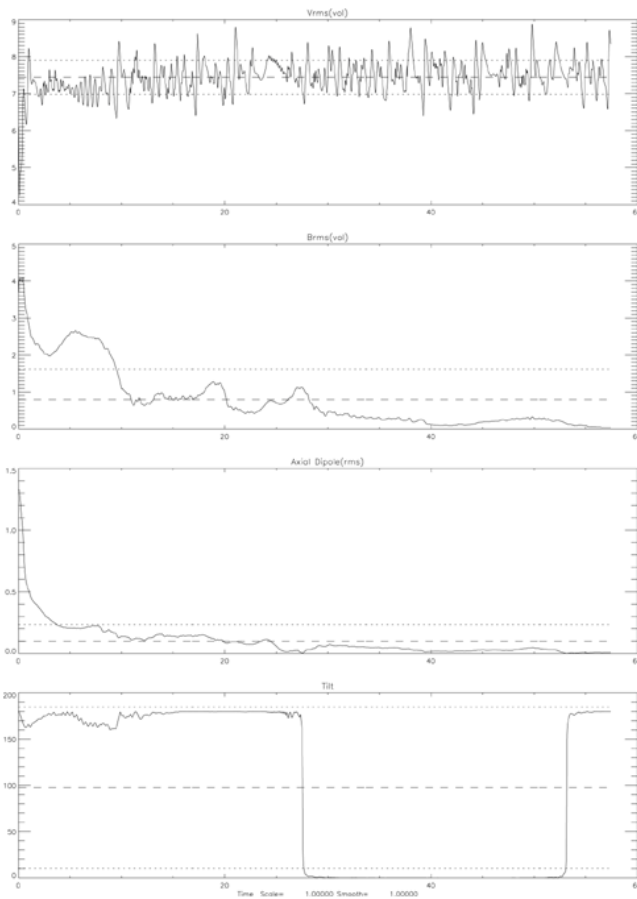
```
$ magx32s4 <par.Rev >p.rev &
```

MAG will then execute and produce datafiles in the same directory. See Appendix B in the MAG User Manual for details on MAG's output files.

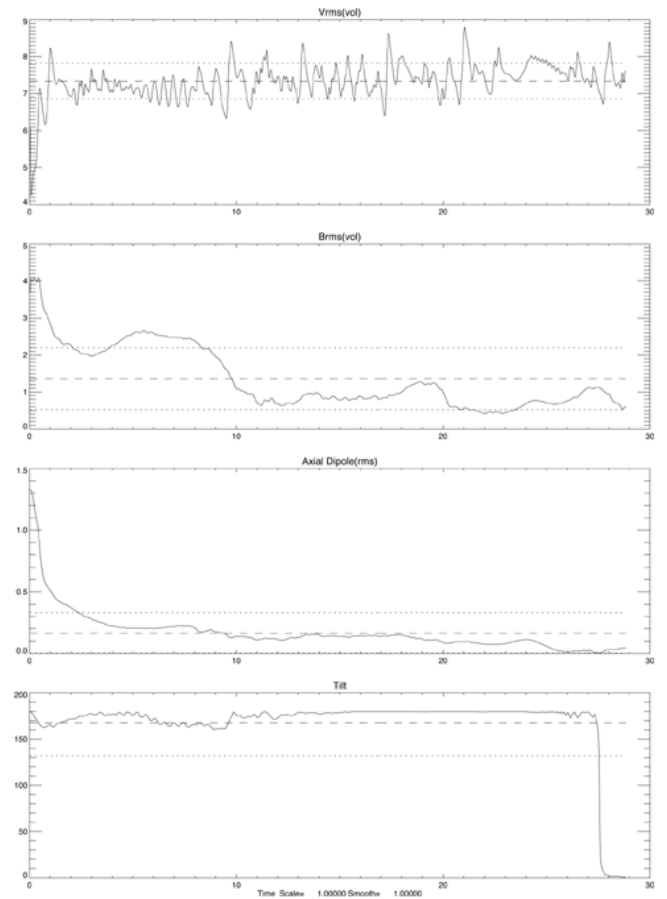
## Results and Discussion

This case has run on 32-bit and 64-bit Intel processors. Figure 1 below shows a plot of mean velocity  $V_{rms}$ , mean magnetic field  $B_{rms}$ , the axial dipole and the dipole tilt on the outer boundary. It indicated a magnetic field reversal between time steps 25 and 30.

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**Figure 1**



**Figure 2**

Figure 2 shows a longer run of MAG, where we see the magnetic field reversed again. At this time, the magnetic field had weakened substantially.

In Figures 3 and 4 (below), the top figure is the pole plot before the second field reversal and the bottom is the pole plot after the second field reversal.

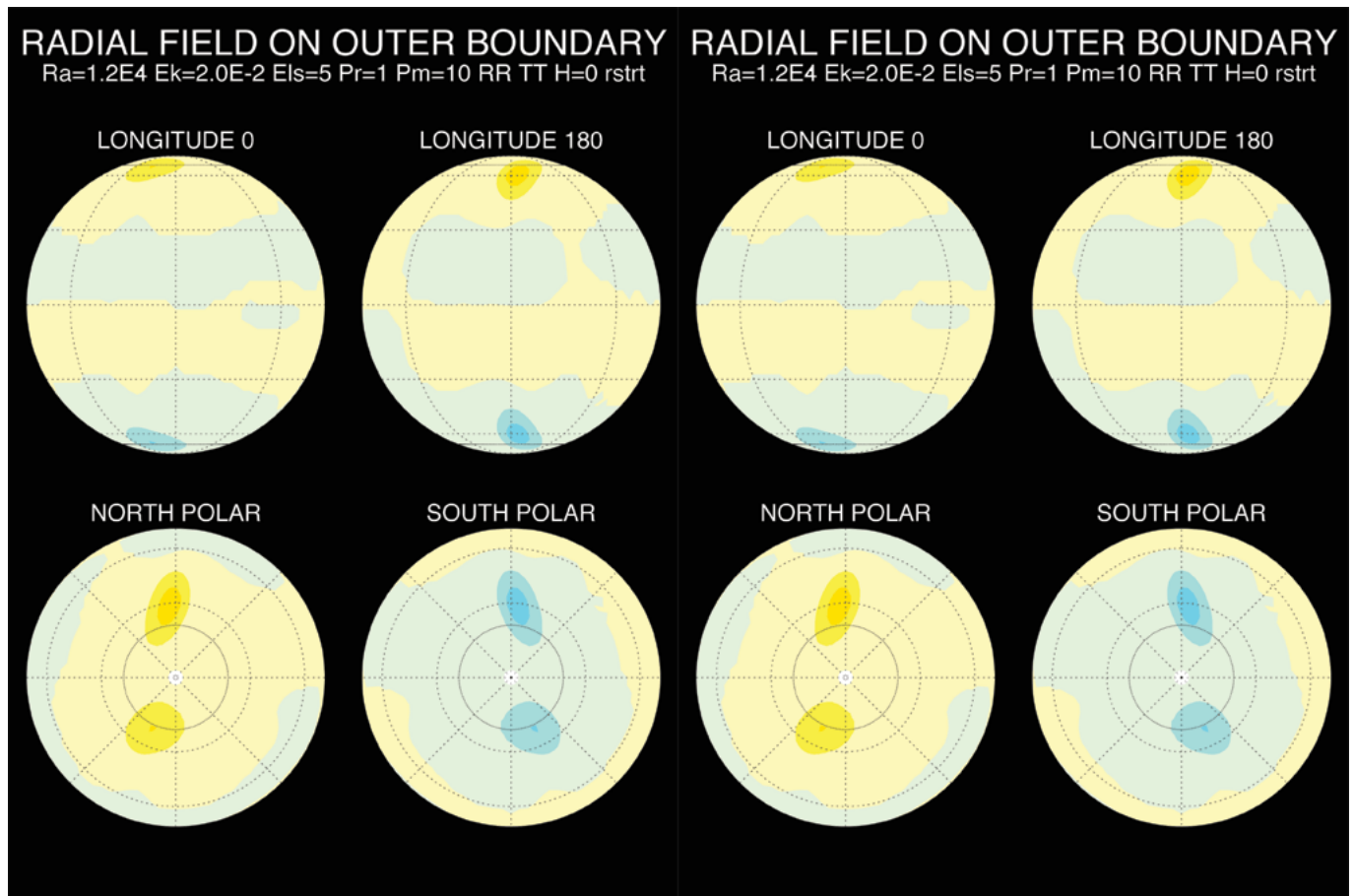


Figure 3

Figure 4

An additional tutorial on creating a Reversal Dynamo movie is found in Chapter 5 of the MAG User Manual.